

# Stars Through Light and Color

*Presented By:*

*Hadar D. / Tzuriya L. / Kim K. / Roni H. / Shon A. / Sally S.*



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# Introduction

- Star- an astronomical object consisting of a luminous spheroid of plasma held together by its own gravity.
- Light- electromagnetic radiation that can be detected by the human eye.



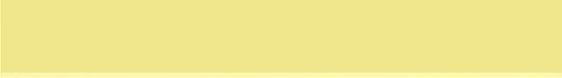
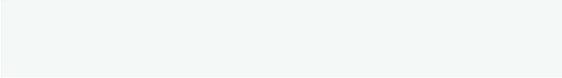
# Observing the Stars

*Google Sky Map*

TEMPRATURE	COLOR	STAR
11000k	White	Rigel
9940k	Blue	Sirius
6530k	Yellow	Procyon
3590k	Red	Betelgeuse
3910k	Orange	Aldebaran
4970k	Yellow	Capella



# First Experiment

Color	Electrical voltage
	0.5v
	0.7v
	0.9v
	1.1v
	1.3v
	1.5v
	1.7v
	1.9v
	2.1v
	2.3v



0.5v



0.7v



0.9v



1.1v



1.3v



1.5v



1.7v



1.9v



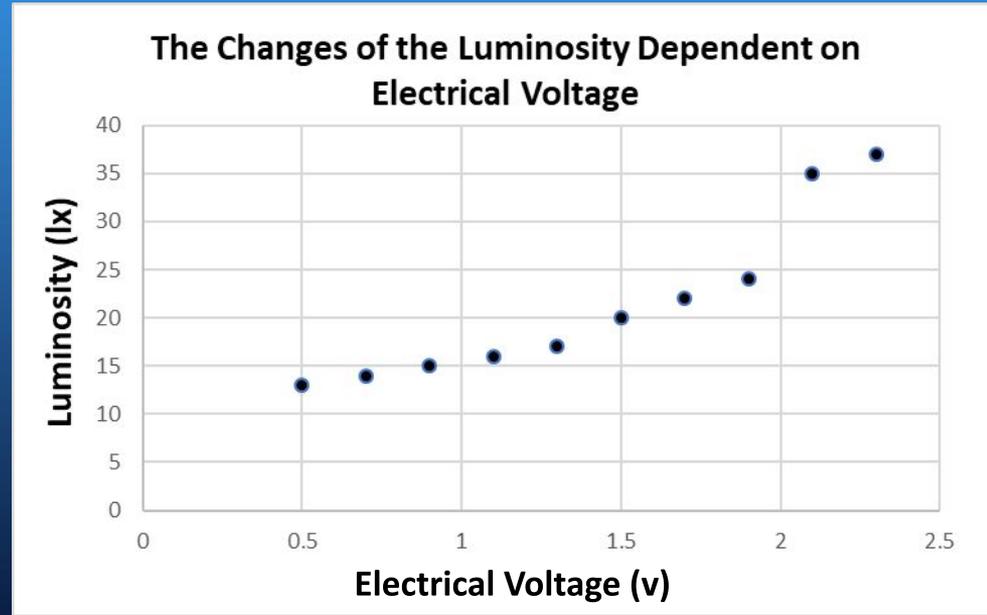
2.1v



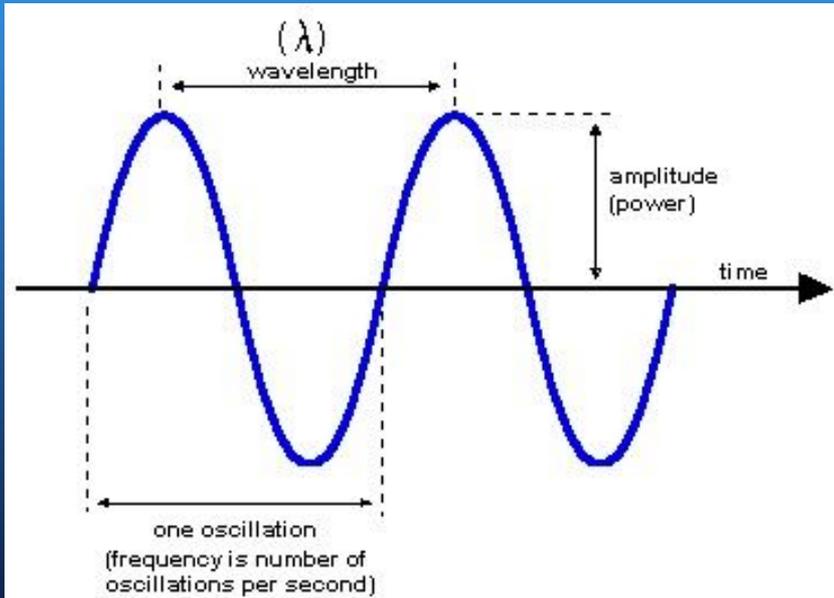
2.3v

# Electrical Voltage Effect on Luminosity

Electrical Voltage	Luminosity
0.5 v	13 lx
0.7 v	14 lx
0.9 v	15 lx
1.1 v	16 lx
1.3 v	17 lx
1.5 v	20 lx
1.7 v	22 lx
1.9 v	24 lx
2.1 v	35 lx
2.3 v	37 lx



# Waves and Wavelength

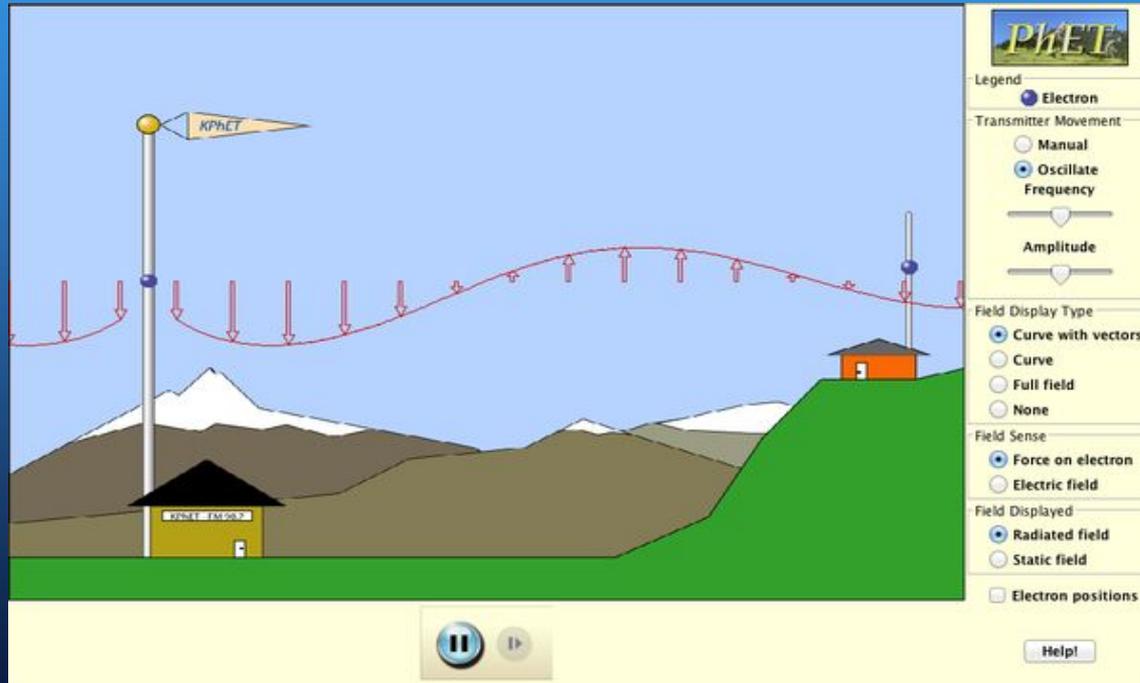


**Wave**- propagation of disturbances from place to place in a regular and organized way

**Amplitude**- the maximum displacement or distance moved by a point on a vibrating body or wave measured from its equilibrium position.

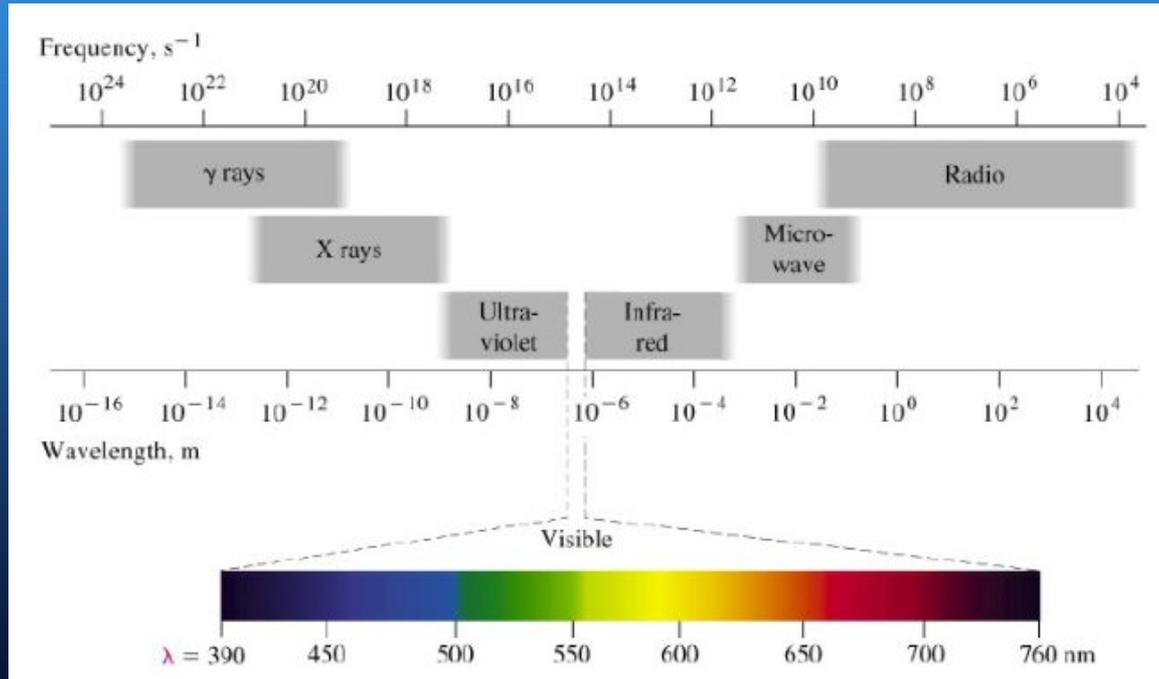
**Wavelength**- distance between corresponding points of two consecutive waves.

# Electromagnetic Waves



□ <https://phet.colorado.edu/en/simulation/legacy/radio-waves>

# Spectrum



**Spectrum**- the arrangement according to wavelength of visible, ultraviolet, and infrared light.

color	wavelength Interval
<a href="#">red</a>	~ 630–700 nm
<a href="#">orange</a>	~ 590–630 nm
<a href="#">yellow</a>	~ 560–590 nm
<a href="#">green</a>	~ 490–560 nm
<a href="#">blue</a>	~ 450–490 nm
<a href="#">violet</a>	~ 400–450 nm

# Second Experiment



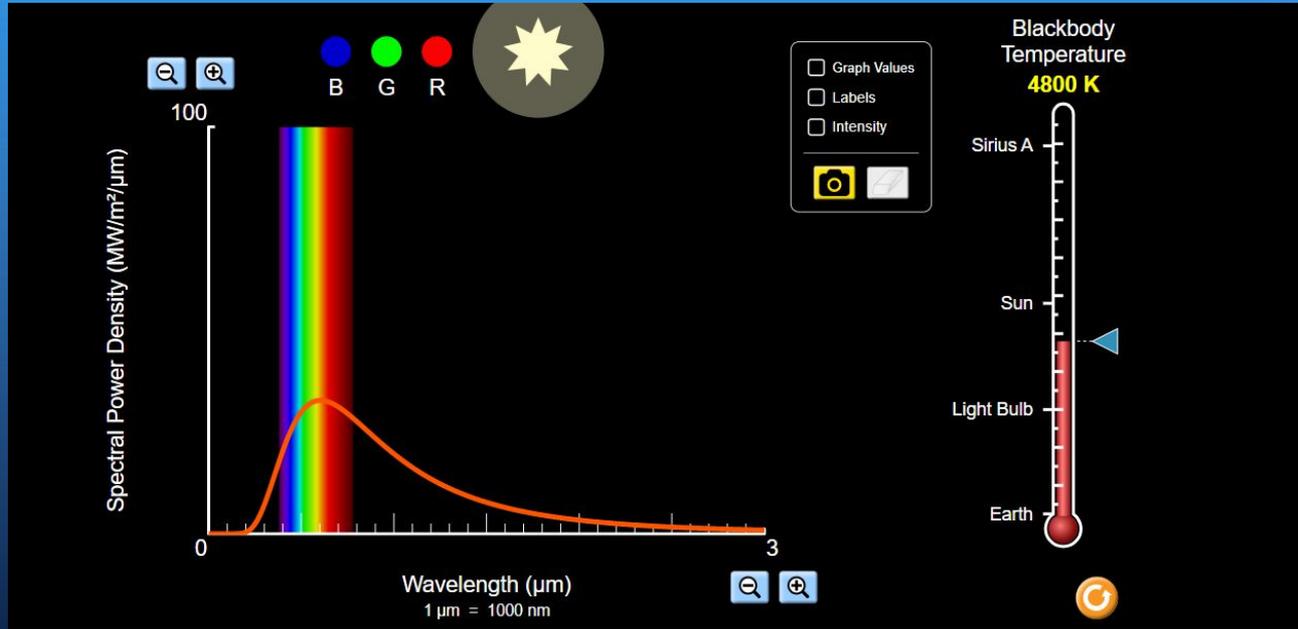
# Second Experiment *(Continue)*



# Second Experiment *(Continue)*



# Blackbody Spectrum



□ [https://phet.colorado.edu/sims/html/blackbody-spectrum/latest/blackbody-spectrum\\_en.html](https://phet.colorado.edu/sims/html/blackbody-spectrum/latest/blackbody-spectrum_en.html)

# Conclusion



The amount of LIGHT produced by an object at EACH WAVELENGTH depends on the TEMPERATURE of the object producing the light.

- ☐ Stars cooler than the sun put out most of their light in the red and infrared regions of the spectrum.
- ☐ Stars hotter than the sun put out most of their light in the blue and ultraviolet regions of the spectrum.



- How to measure a wavelength by using a spectrometer?
- How are wavelength and temperature related?